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### Type I X-Ray Bursts: Populations, Rare Specimen, and Disk Interaction

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- Much effort in improving nuclear reaction data of rare p-rich isotopes
- Predictions of burst regimes still do not match observations
- ► NS EOS: atmosphere models used, but uncertain effect of disk



## WFM and LAD view of Bursts and Superbursts



Keek, Heger, In 't Zand 2012; In 't Zand, ..., Keek, et al. 2015

- Superburst quenches regular bursts for a day
- Transition from stable H/He burning to mHz QPOs to X-ray bursts









- MINBAR: all RXTE/PCA and BeppoSAX/WFC bursts
- Short recurrence times from H-accretors: new models from Keek & Heger 2017
- Reduced burst rate above ~ 20% Eddington: WFM opportunity







In 't Zand, Mahmoodifar, Keek, et al. 2015

- Ignition near the pole (left) and at the equator (right)
- Track the phase of pulsations to follow flame spreading
- Low-amplitude pulsations as seen in 4U 1636-536 2001 superburst?

# Superbursts and Long Helium Bursts





Keek, Cumming, et al. 2016

In 't Zand & Weinberg 2014

- Slewing on WFM trigger
- First 1000 s of superburst constrain carbon flame propagation
- First few minutes of intermediate duration burst show superexpansion and absorption edges from NS surface: EOS constraint







- Push out inner disk or infall from radiation drag?
- Heating and ionization of the disk; launching of winds?
- Coronal cooling; boundary layer?
- Reflection spectroscopy provides a diagnostic



## Model X-Ray Reflection Spectra





Ballantyne 2004

- $\blacktriangleright$  6.4 keV Iron line depends on ionization parameter  $\xi$
- Soft band depends on density and composition







Strohmayer & Brown 2002

Keek, Ballantyne et al. 2016

- Superburst 4U 1820-30 (*RXTE*/PCA): iron line+edge
- Helium burst IGR J17062-6143 (Swift/XRT): soft excess







- Seen 2x with Chandra and Swift: present in 3 NICER bursts
- Signature of burst impact on environment: enhanced accretion or reflection?
- Increased accretion by  $f_{\rm a} \simeq 2.7$ , or a reflection fraction of  $\sim 0.5$







Keek, Wolf, and Ballantyne 2016

- ► 15×NICER+LOFT: very detailed light curves and spectra
- In 10 s, measure log ξ to 0.12%, R<sub>in</sub> to 14%, disk inclination to 12%, emissivity profile to 4%





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#### WFM

- Large sample of bursts
- Better constraints on burning regimes, such as transition to stable burning (carbon creation for superbursts)
- Triggers for rare long bursts (superbursts, intermediate duration bursts)
- Timing with XRCA and LAD
  - Track phase of burst oscillations
  - Measure spin of bursting neutron stars
  - Low-amplitude pulsations as seen in superburst
- XRCA and LAD spectroscopy
  - Burst impact on accretion environment
  - Detailed reflection features; accretion disk tomography
  - Evolution of persistent spectral component: inner disk, spreading layer, corona, etc.